Levinson theorem for Dirac particles in one dimension

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Abstract. The scattering of Dirac particles by symmetric potentials in one dimension is studied. A Levinson theorem is established. By this theorem, the number of bound states with even(odd)-parity, n_{+} (n−), is related to the phase shifts $\eta_+(\pm E_k)$ [$\eta_-(\pm E_k)$] of scattering states with the same parity at zero momentum as follows:

$$
\eta_{\pm}(\mu) + \eta_{\pm}(-\mu) \pm \frac{\pi}{2} [\sin^2 \eta_{\pm}(\mu) - \sin^2 \eta_{\pm}(-\mu)] = n_{\pm} \pi.
$$

The theorem is verified by several simple examples.

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1 Introduction

In 1949, Levinson established a theorem in nonrelativistic quantum mechanics [1]. The theorem gives a relation between bound states and scattering states in a given angular momentum channel l, i.e., the total number of bound states n_l is related to the phase shift $\delta_l(k)$ at threshold $(k = 0)$:

$$
\delta_l(0) = n_l \pi, \quad l = 0, 1, 2, \dots \tag{1a}
$$

The case $l = 0$ should be modified as

$$
\delta_0(0) = (n_0 + 1/2)\pi
$$
 (1b)

when there exists a zero-energy resonance (a half-bound state) [2]. This is one of the most interesting and beautiful results in nonrelativistic quantum theory. The subject was then studied by many authors (some are listed in the Refs. [2–8]) and generalized to relativistic quantum mechanics $[6,9-15]$. In reference $[6]$ the theorem was first written as

$$
\delta_l(0) - (-)^l \frac{\pi}{2} \sin^2 \delta_l(0) = n_l \pi.
$$
 (2)

The second term on the left-hand side (lhs) appears automatically in the Green function approach to Levinson's theorem, which was used in earlier works [3,4] and further developed by Ni [6]. This method is quite different from the original one where the theorem is obtained by using the analytic properties of the Jost functions, and is con-

venient for generalization to the relativistic case. It can be shown that $\delta_l(0)/\pi$ takes integers when $l > 0$ or when $l = 0$ but there exists no zero-energy solution. Then equation (2) is equivalent to equation (1a). When $l = 0$ and there exists a zero-energy solution (a half-bound state), however, $\delta_l(0)/\pi$ takes half integers and $\sin^2 \delta_0(0) = 1$. In this case equation (2) is equivalent to equation (1b). Thus equation (2) combines the two cases in equation (1) into a unified form. The correct generalization of Levinson's theorem to Dirac particles was first obtained by the Green function method in reference [10]. If the phase shifts in the angular momentum channel κ ($\kappa = \pm 1, \pm 2, \ldots$) are denoted by $\eta_{\kappa}(\pm E_k)$ and the total number of bound states in that channel by n_{κ} , then the Levinson theorem read

$$
\eta_{\kappa}(\mu) + \eta_{\kappa}(-\mu) - (-)^{\kappa} \epsilon(\kappa) \frac{\pi}{2} [\sin^2 \eta_{\kappa}(\mu) - \sin^2 \eta_{\kappa}(-\mu)] = n_{\kappa} \pi,
$$
\n(3)

where $\epsilon(\kappa)=1$ (−1) for positive (negative) κ , μ is the mass of the particle. The effect of half-bound states has been automatically included in this equation, though the problem is more complicated than in the nonrelativistic case. There is a modulo- π ambiguity in the definition of the phase shifts. This is common to both nonrelativistic and relativistic cases. In the former case this was resolved by setting $\delta_l(\infty) = 0$ (rather than a multiple of π) which can be freely done. In the latter case one is not allowed to set $\eta_{\kappa}(\pm\infty) = 0$, but the modulo- π ambiguity can be appropriately resolved [9,10,16,17].

Most of the above cited works mainly deal with the problem in ordinary three-dimensional space. Recently, several authors have studied the two-dimensional version

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of Levinson's theorem for both nonrelativistic Schrödinger particles and relativistic Dirac particles [18–20]. In the works of the present author the Green function method is employed [18]. In the nonrelativistic case we obtain

$$
\delta_m(0) = n_m \pi, \quad m = 0, 1, 2, \dots \tag{4}
$$

In two dimensions $\delta_m(0)/\pi$ always take integers. This threshold behavior is quite different from that in three dimensions and in one dimension (see below). In reference [19], Dong, Hou, and Ma studied the problem by the method of Sturm-Liouville theorem and found that the case $m = 1$ in equation (4) should be modified as

$$
\delta_1(0) = (n_1 + 1)\pi
$$
 (5)

when there exists a half-bound state in this channel. In other words, the half-bound state with $m = 1$ plays the role of a real bound state in the Levinson theorem. The same result was also obtained in an earlier paper [20]. Similar modifications also occur in the relativistic case. The miscounting in our work may be due to the very subtle behavior of the above state and the shortcomings of the regularization procedure we employed (cf. Eqs. (44–47) in this paper), rather than the failure of the completeness relation as remarked in reference [19] (see some more discussions in Sect. 2). The above state is quite different from the half-bound state with $m = 0$ and from the halfbound states in three and one dimension because it tends to zero at infinity. As a result it is lost in the regularization procedure. In three- and one-dimensional space there is no similar case and the Green function method leads to correct results.

It seems that the one-dimensional version of Levinson's theorem has attracted more attention than the two-dimensional one. In fact, the nonrelativistic case has been studied by several authors [21–23]. In a symmetric potential $V(x)$ (an even function of x), the Levinson theorem takes the form

$$
\delta_{\pm}(0) \pm \frac{\pi}{2} \sin^2 \delta_{\pm}(0) = n_{\pm} \pi,
$$
\n(6)

where $+ (-)$ indicates even(odd)-parity. It is easily seen that the odd-parity case coincides with the case $l = 0$ in three dimensions, while the even-parity case has no counterpart in three dimensions. This is the main feature in one dimension.

The purpose of this paper is to generalize the onedimensional Levinson theorem to relativistic Dirac particles. The interest of this problem is threefold. First, the Levinson theorem is a nonperturbative result. It may be useful in the study of nonperturbative field theories. In view of the wide interest in field theories and condensed matter physics in lower dimensions in recent years, and the potential applications of the Levinson theorem, the problem seems of interest. Indeed, some applications of the theorem to field theories have appeared in the literature [21,24,25]. More recently, the Levinson theorem in two dimensions has been used [26] to the study of the screened Coulomb potential which plays an important role

in the physics of semiconductor heterostructures. Second, as an academic problem, it exhibits some new features as compared with the three- or two-dimensional problem. Thus it may be of some interest in itself, as many onedimensional models, say, the one-dimensional QED (the Schwinger model), studied in the literature. Third, it may be related to the real three-dimensional problem. Consider a Dirac particle in an external potential V (the zero component of a vector potential). If V depends only on one space variable, say, x, one may consider a special case where the particle moves only in the x direction. This special case may be described by the one-dimensional Dirac equation. This connects the problem to the real physical world.

Consider Dirac particles in a symmetric potential in one dimension. The phase shifts of even(odd)-parity solutions (see Sect. 2) are denoted by $\eta_+(\pm E_k)$ $[\eta_-(\pm E_k)]$, while the number of bound states with the same parity is denoted by n_{+} (n−). The Levinson theorem connects them with each other:

$$
\eta_{\pm}(\mu) + \eta_{\pm}(-\mu) \pm \frac{\pi}{2} [\sin^2 \eta_{\pm}(\mu) - \sin^2 \eta_{\pm}(-\mu)] = n_{\pm} \pi.
$$
\n(7)

We will establish this theorem for cutoff potentials (vanish when $|x| > a > 0$ in this paper. Though the even(odd)parity case coincides in appearance with the case $\kappa = 1$ $(\kappa = -1)$ in three dimensions (*cf.* Eq. (3)). They are essentially different. First, the one-dimensional equation is different from the radial one with $\kappa = 1$ or $\kappa = -1$ in three dimensions. The boundary conditions at the origin are also different. Second, the two critical energy solutions with $E = \pm \mu$ and even(odd)-parity, if exist, are both half-bound states in one dimension, while one of the two in three dimensions with $\kappa = 1$ ($\kappa = -1$) is a bound state. Third, the threshold behaviors of the phase shifts are different. For example, in three dimensions $\eta_1(\mu)/\pi$ always takes integers, while in one dimension $\eta_+(\mu)/\pi$ takes integers only when there exists a half-bound state with $E = \mu$ and even parity. In general $\eta_{+}(\mu)/\pi$ takes half integers. There is a similar difference between $\eta_-(-\mu)$ and $\eta_{-1}(-\mu)$. Therefore each case in equation (7) has no counterpart in three dimensions.

Throughout this paper we use natural units where $\hbar = c = 1$. In the next section we first discuss the solutions of the one-dimensional Dirac equation in an external symmetric potential. Then we discuss the scattering problem. In Section 3 the behavior of the phase shifts near $k = 0$ is analyzed. The behavior when there exists a half-bound state is distinguished from the case without one. In Section 4 we establish the Levinson theorem for cutoff potentials. In this paper we do not resort to the Green function method developed in reference [6] and used in references [10,17,18]. We employ the completeness of the solutions of the Dirac equation to derive the theorem more directly. This is essentially equivalent to the Green function method but simpler, and has been used in some recent works [13,23]. In Section 5, we first discuss the behavior of the phase shifts at infinite momentum and the resolution of the modulo- π ambiguity in the definition of the phase shifts. Then we examine the Levinson theorem by two simple examples. Finally we give a brief summary of the results.

2 Dirac particles in one dimension

We work in $(1+1)$ -dimensional space-time. The Dirac equation in an external vector field $A_{\nu}(t, x)$ reads

$$
(\mathrm{i}\gamma^{\nu}D_{\nu}-\mu)\Psi=0,\qquad(8)
$$

where μ is the mass of the particle, $D_{\nu} = \partial_{\nu} + ieA_{\nu}$, e is the coupling constant, and summation over the repeated Greek index ν ($\nu = 0, 1$) is implied. The γ^{μ} are Dirac matrices satisfying the Clifford algebra:

$$
\{\gamma^{\mu}, \gamma^{\nu}\} = 2g^{\mu\nu},\tag{9}
$$

where $g^{\mu\nu} = \text{diag}(1, -1)$ is the Minkowskian metric. In this paper we only consider the zero component of A_{ν} , which is an even function of x , namely, we consider the special case where

$$
A_1 = 0, \quad eA_0 = V(x), \tag{10}
$$

where $V(x)$ is symmetric with respect to reflection:

$$
V(-x) = V(x). \tag{11}
$$

In this case we may set

$$
\Psi(t,x) = e^{-iEt}\psi(x),\tag{12}
$$

and get a stationary equation for $\psi(x)$:

$$
H\psi = E\psi,\tag{13}
$$

where the Hamiltonian

$$
H = \alpha p + \gamma^0 \mu + V(x),\tag{14}
$$

where $p = -i\partial_r$ is the momentun operator and $\alpha = \gamma^0 \gamma^1$.

To solve the Dirac equation (13) an explicit representation of the Dirac matrices is necessary. In one dimension this can be realized by the Pauli matrices:

$$
\gamma^0 = \sigma^3, \quad \gamma^1 = i\sigma^1. \tag{15}
$$

In this representation $\alpha = -\sigma^2$. We denote the twocomponent spinor $\psi(x)$ as

$$
\psi(x) = \begin{pmatrix} u(x) \\ v(x) \end{pmatrix} = \begin{pmatrix} u_1(x) \\ u_2(x) \end{pmatrix}, \tag{16}
$$

where the second notation will be used only in a few occasions where summation over the spinor index is involved, then equation (13) can be explicitly written as a system of first-order differential equations for u and v :

$$
u' + (E + \mu - V)v = 0, \quad v' - (E - \mu - V)u = 0. \tag{17}
$$

Since $V(x)$ is an even function of x, it is easy to show that if $(u(x), v(x))^{\tau}$ is a solution of equation (17), then $(u(-x), -v(-x))^{\tau}$ is also a solution with the same energy E, where τ denotes matrix transposition. Thus the solutions of equation (17) can be chosen to have definite parities. Even-parity solutions are denoted by $\psi(x,+)$ and defined by the property of its components under reflection:

$$
u(-x, +) = u(x, +), v(-x, +) = -v(x, +).
$$
 (18a)

Odd-parity solutions are denoted by $\psi(x, -)$ and defined by

$$
u(-x,-) = -u(x,-), \quad v(-x,-) = v(x,-). \tag{18b}
$$

If $(u(x), v(x))^{\tau}$ and $(\tilde{u}(x), \tilde{v}(x))^{\tau}$ are both solutions to equation (17) with the same energy value E , it can be shown that $u\tilde{v} - \tilde{u}v = constant$. For bound states, all the functions u, v, \tilde{u} , and \tilde{v} must vanish at infinity, so the constant in the above equation is zero, and we have $\tilde{u}/u = \tilde{v}/v$. We denote this fraction by w, a function of x, and get $\tilde{u} = wu$, $\tilde{v} = wv$. Substituting this into equation (17), we have $w'u = 0$, $w'v = 0$. As a nontrivial solution, u and v cannot simultaneously vanish at any point, so we have $w' = 0$, or $w =$ constant, which means that the two solutions are equivalent. Thus there is no degeneracy with bound states.

From the above discussions we conclude that a bound state solution of equation (17) must have definite parity. On the other hand, scattering states need not vanish at infinity. For a given energy, there are two linearly independent solutions. Both are physically acceptable. They do not necessarily have definite parities. However, they can be chosen such that one is of even parity and the other of odd. The above discussion is applicable to the case of free particles where $V = 0$ and there is no bound state.

For free particles, $V = 0$, then equation (17) can be easily solved. We have positive-energy solutions with $E > \mu$ and negative-energy solutions with $E < -\mu$, both being scattering solutions. We define $k = \sqrt{E^2 - \mu^2} \ge 0$, and denote positive(negative)-energy solutions by the subscript k (−k), thus we have, say,

$$
E_{\pm k} = \pm E_k = \pm \sqrt{k^2 + \mu^2}.
$$
 (19)

The solutions will be given below. It is remarkable that when $E = \mu$ there is a nontrivial solution (not identically zero) with even-parity while when $E = -\mu$ there is one with odd-parity. These are usually called half-bound states.

Now we consider particles moving in the external symmetric potential $V(x)$. In this section the potential need not be a cutoff one. We assume that $V(x)$ decreases more rapidly than x^{-2} when $x \to \infty$, is less singular than x^{-1} when $x \to 0$, and is regular everywhere (except possibly at $x = 0$). It can be shown that $E^2 > \mu^2$ gives scattering solutions while $E^2 < \mu^2$ gives bound state solutions. (The critical energy states with $E = \pm \mu$ will be discussed in the next section.) Scattering states will be denoted as above.

Bound states with even(odd)-parity will be denoted by a subscript κ_+ (κ_-) which takes discrete values. Scattering states are normalized by Dirac δ functions. These are not necessary for the following discussions and are thus omitted. The orthonormal relation for bound states is given by

$$
(\psi_{\kappa'_{\pm}}, \psi_{\kappa_{\pm}}) = \int_{-\infty}^{+\infty} dx \, \psi_{\kappa'_{\pm}}^{\dagger}(x) \psi_{\kappa_{\pm}}(x) = \delta_{\kappa'_{\pm} \kappa_{\pm}}, \quad (20)
$$

which will be used below. Since the Hamiltonian (14) is an Hermitian operator, we have the completeness relation

$$
\int_{0}^{\infty} dk \left[\psi_{k}(x, +) \psi_{k}^{\dagger}(x', +) + \psi_{k}(x, -) \psi_{k}^{\dagger}(x', -) \right. \n+ \psi_{-k}(x, +) \psi_{-k}^{\dagger}(x', +) + \psi_{-k}(x, -) \psi_{-k}^{\dagger}(x', -) \n+ \sum_{\kappa_{+}} \psi_{\kappa_{+}}(x) \psi_{\kappa_{+}}^{\dagger}(x') + \sum_{\kappa_{-}} \psi_{\kappa_{-}}(x) \psi_{\kappa_{-}}^{\dagger}(x') = \delta(x - x').
$$
\n(21)

It should be remarked that all regular solutions should be included in the summation or integration. Thus halfbound states, when they appear, should be included, or the above equation does not hold. This can be verified straightforwardly in the free case where two half-bound states are present as pointed out above. The asymptotic forms of the various types of scattering solutions are given by

$$
u_{\pm k}(x, +) \rightarrow \pm \sqrt{\frac{E_k \pm \mu}{2\pi E_k}} \cos[k|x| + \eta_+(\pm E_k)], \qquad (22a)
$$

$$
v_{\pm k}(x,+) \to \sqrt{\frac{E_k \mp \mu}{2\pi E_k}} \epsilon(x) \sin[k|x| + \eta_+(\pm E_k)], \quad (22b)
$$

$$
u_{\pm k}(x,-) \to \sqrt{\frac{E_k \pm \mu}{2\pi E_k}} \epsilon(x) \sin[k|x| + \eta_{-}(\pm E_k)], \quad (22c)
$$

$$
v_{\pm k}(x,-) \to \mp \sqrt{\frac{E_k \mp \mu}{2\pi E_k}} \cos[k|x| + \eta_{-}(\pm E_k)], \qquad (22d)
$$

when $x \to \infty$ (+ ∞ or $-\infty$), where $\epsilon(x)=1$ (-1) for positive (negative) x, $\eta_+(\pm E_k)$ [$\eta_-(\pm E_k)$] are the phase shifts of even(odd)-parity solutions. The phase shifts depend on the sign as well as the magnitude of the energy, as equation (17) does. By setting all the phase shifts to zero, the right-hand side (rhs) of equation (22) gives exactly the solutions for free particles. Compared with the free solutions, the asymptotic forms in the external symmetric potential are distorted by the phase shifts. But it should be remarked that the normalization factors in equation (22) are the same as in the free case.

It is well known that positive(negative)-energy solutions correspond to particles (antiparticles) after second quantization. In the following we give some results for the scattering of positive-energy solutions by the symmetric potential $V(x)$ described above. The results for the scattering of negative-energy solutions are similar.

We denote the reflection amplitude by R_k and the transmission one by T_k . As a consequence of charge conservation, they satisfy

$$
|R_k|^2 + |T_k|^2 = 1.
$$
 (23)

Similar to the partial-wave method in three or two dimensions, R_k and T_k can be expressed in terms of the phase shifts. The result reads

$$
R_k = \frac{1}{2} [e^{2i\eta_+(E_k)} - e^{2i\eta_-(E_k)}]
$$

= ie^{i[$\eta_+(E_k)+\eta_-(E_k)$] sin[$\eta_+(E_k) - \eta_-(E_k)$], (24a)}

$$
T_k = \frac{1}{2} [e^{2i\eta_+(E_k)} + e^{2i\eta_-(E_k)}]
$$

= $e^{i[\eta_+(E_k) + \eta_-(E_k)]} \cos[\eta_+(E_k) - \eta_-(E_k)].$ (24b)

These results obviously satisfy equation (23). They hold for both left-incident particles and right-incident ones since the potential is symmetric. If the potential is not symmetric, the reflection amplitude would depend on the direction of incidence. We can construct a scattering matrix and discuss various properties of it as in the nonrelativitic case [22]. However, we will not go further in this respect. In the following we will confine our discussions to symmetric potentials.

From the above discussions, we see that information about the scattering process is contained in the phase shifts. The latter are determined by solving the system of equations (17) with the boundary conditions (22), and thus depend on the particular form of the potential. On the other hand, the total number of bound states with definite parities also depends on the particular form of the potential. The purpose of the Levinson theorem is to establish a relation between the phase shifts and the total number of bound states with a given parity.

3 Phase shifts near threshold

In this section we discuss the behavior of the phase shifts $\eta_{\pm}(E_k)$ and $\eta_{\pm}(-E_k)$ near $k=0$. This is important in the development of the Levinson theorem, and is also helpful in understanding the result.

From now on we consider cutoff potentials that satisfy $V(x) = 0$ when $|x| > a > 0$. Such potentials will be denoted by $V_a(x)$ in the following. Since we will deal with solutions with definite parities, we need only consider the region $x \geq 0$. The solutions in the region $x > a$ will be indicated by a superscript >. Scattering solutions are given by those on the rhs of equation (22) since $V_a(x) = 0$ in this region. Bound state solutions will be discussed below. In the region $x < a$, the solutions are not explicitly available. First we consider the behavior of the solutions near $x = 0$ and set appropriate boundary conditions. As assumed in Section 2, $V_a(x)$ is less singular than x^{-1} when $x \to 0$. If $V_a(x)$ is regular at $x = 0$, the appropriate boundary conditions for equation (17) are obviously

where

$$
u(0,+) = 1, \quad v(0,+) = 0,
$$
 (25a)

$$
u(0, -) = 0, \quad v(0, -) = 1. \tag{25b}
$$

If $V(x)$ behaves like $A_0/x^{1-\rho/2}$ when $x \to 0^+$ where A_0 is a constant and $0 < \rho < 2$, it can be shown that the above boundary conditions remain valid. These boundary conditions are applicable to both scattering solutions and bound state solutions, with positive or negative energy values. Let us consider scattering solutions of equation (17) with positive energy E_k that satisfy the boundary conditions (25). Note that equation (17) depends on k^2 rather than k , and the boundary conditions (25) are independent of k. Thus these solutions should only depend on k^2 as well. In other words, they are even functions of k . We will denote them by $f_{\pm}(x, k^2)$ and $g_{\pm}(x, k^2)$, where the subscript + $(-)$ indicates even(odd)-parity. Thus scattering solutions with positive energy E_k are given in the region $x < a$ by

$$
u_k^{\leq}(x, \pm) = A_{\pm}(k) f_{\pm}(x, k^2),
$$

$$
v_k^{\leq}(x, \pm) = A_{\pm}(k) g_{\pm}(x, k^2),
$$
 (26)

where the superscript \lt indicates the region $x \lt a, A_{+}(k)$ are constants which in general depend on k such that the solutions in the two regions can be appropriately connected. Obviously, $u_k(x, \pm)$ and $v_k(x, \pm)$ should be continuous at $x = a$, so that the probability density and the probability current density are continuous at the point. The phase shifts $\eta_{\pm}(E_k)$ are determined by this condition. The results are given by

$$
\tan \eta_{+}(E_{k}) = \frac{\beta_{+}(\xi) - \tan \xi}{1 + \beta_{+}(\xi) \tan \xi},
$$

$$
\tan \eta_{-}(E_{k}) = \frac{1 + \beta_{-}(\xi) \tan \xi}{\tan \xi - \beta_{-}(\xi)},
$$
(27a)

where $\xi = ka$ and $\beta_{\pm}(\xi)$ are defined by

$$
\beta_{\pm}(\xi) = \sqrt{\frac{E_k + \mu}{E_k - \mu}} \frac{g_{\pm}(a, k^2)}{f_{\pm}(a, k^2)}.
$$
 (27b)

The above results show that the behavior of $\eta_+(E_k)$ is determined by that of $\beta_{\pm}(\xi)$ and ultimately by that of $f_{\pm}(a, k^2)$ and $g_{\pm}(a, k^2)$. The general dependence of $f_{\pm}(x, k^2)$ and $g_{\pm}(x, k^2)$ on k may be very complicated since equation (17) depends on k in a rather complicated way. Fortunately, only the property of $f_{\pm}(x, k^2)$ and $g_{\pm}(x, k^2)$ near $k = 0$ is necessary for our purpose.

We consider the limit $E = E_k \rightarrow \mu$ $(k \rightarrow 0)$ of equation (17). In this limit it takes the following form to the first order in k^2 :

$$
u'(x) + P(x, k2)v(x) = 0,v'(x) + Q(x, k2)u(x) = 0,
$$
\n(28a)

 $P(x, k^2) = 2\mu + \frac{k^2}{2}$ $\frac{\kappa}{2\mu}-V(x),$ $Q(x, k^2) = -\frac{k^2}{2\mu} + V(x).$ (28b)

The equations depend only on k^2 . Their solutions that satisfy the boundary conditions (25) should depend only on k^2 as well, since the boundary conditions do not depend on k. These solutions will be denoted by $\tilde{f}_{\pm}(x, k^2)$ and $\tilde{g}_{\pm}(x, k^2)$. Note that both $P(x, k^2)$ and $Q(x, k^2)$ are integral functions of k . Thus a theorem of Poincaré tells us that $\tilde{f}_{\pm}(x, k^2)$ and $\tilde{g}_{\pm}(x, k^2)$, which satisfy k independent boundary conditions, are also integral functions of k. On the other hand, equation (17) coincides with equation (28) in the limit $E = E_k \rightarrow \mu \ (k \rightarrow 0)$. Therefore, $f_{\pm}(x, k^2)$ and $g_{\pm}(x, k^2)$ must coincide with $\tilde{f}_{\pm}(x, k^2)$ and $\tilde{g}_{\pm}(x, k^2)$ respectively in the limit $k \to 0$, since they satisfy the same boundary conditions. Hence we conclude that $f_{\pm}(x, k^2)$ and $g_{\pm}(x, k^2)$ are analytic functions of k in the neighbourhood of $k = 0$. The above analysis holds regardless of whether the potential is cutoff or not, and the functions $f_{\pm}(x, k^2)$ and $g_{\pm}(x, k^2)$ are analytic in k near $k = 0$ for any fixed x in the interval $[0, +\infty)$, not only in $[0, a]$.

We have shown that $f_{\pm}(a, k^2)$ and $g_{\pm}(a, k^2)$ are even functions of k and analytic near $k = 0$. Therefore, when $k \to 0$, the leading term(s) for $g_{\pm}(a, k^2)/f_{\pm}(a, k^2)$ must be given by one of the following forms:

$$
\frac{g_{\pm}(a,k^2)}{f_{\pm}(a,k^2)} \to \alpha_1^{\pm} \xi^{2l_1^{\pm}}, \ \alpha_2^{\pm} \xi^{-2l_2^{\pm}}, \ \alpha_4^{\pm} + \alpha_3^{\pm} \xi^{2l_3^{\pm}}, \ (\xi \to 0)
$$
\n(29)

where l_1^{\pm} , l_2^{\pm} , and l_3^{\pm} are natural numbers, α_1^{\pm} , α_2^{\pm} , α_3^{\pm} and α_4^{\pm} are nonzero constants. Consequently, the leading term for $\beta_{\pm}(\xi)$ is given by one of the following forms:

$$
\beta_{\pm}(\xi) \to \alpha_{\pm} \xi^{2l_{\pm} - 1}, \quad \tilde{\alpha}_{\pm} \xi^{-(2\tilde{l}_{\pm} - 1)}, \quad (\xi \to 0), \quad (30)
$$

where l_{\pm} and l_{\pm} are natural numbers, α_{\pm} and $\tilde{\alpha}_{\pm}$ are nonzero constants. Substituting this into equation (27a) we have

$$
\tan \eta_{+}(E_k) \to b_{+} \xi^{2p_{+}-1} \text{ or } \tilde{b}_{+} \xi^{-(2\tilde{p}_{+}-1)}, \ (\xi \to 0),
$$
\n(31a)

$$
\tan \eta_{-}(E_k) \to b_{-} \xi^{-(2p_{-}-1)} \text{ or } \tilde{b}_{-} \xi^{2\tilde{p}_{-}-1}, \ (\xi \to 0),
$$
\n(31b)

where p_{\pm} and \tilde{p}_{\pm} are natural numbers, b_{\pm} and b_{\pm} are nonzero constants. Thus $\tan \eta_{\pm}(E_k) \to 0$ or ∞ in the limit $k \to 0$ or $\xi \to 0$, and $\eta_{\pm}(\mu)/\pi$ take integers or half integers.

In order to study the threshold behaviour of $\eta_{\pm}(E_k)$ more specifically, we consider bound state solutions of equation (17) with positive energy

$$
E = E_{\lambda} = \sqrt{\mu^2 - \lambda^2}, \quad 0 \le \lambda \le \mu. \tag{32}
$$

The solutions satisfying the boundary conditions (25) in the region $x < a$ will be denoted by $F_{\pm}(x, \lambda^2)$ and $G_{\pm}(x,\lambda^2)$. They are even functions of λ as implied by the notations, since the equations are invariant under the change $\lambda \to -\lambda$ and the boundary conditions are independent of λ . Note that when $-\lambda^2$ is replaced by k^2 , E_λ becomes E_k and the solutions become the above scattering ones since the boundary conditions are the same. Thus we have the useful relations

$$
F_{\pm}(x, -k^2) = f_{\pm}(x, k^2), \quad G_{\pm}(x, -k^2) = g_{\pm}(x, k^2).
$$
 (33)

In the region $x>a$, the solutions are explicitly available since $V_a(x) = 0$. They can be continuously connected to the interior solutions (for $x < a$) only for some specific values of λ , which determine the discrete energy eigenvalues of bound states. This is quite different from the case of scattering states, where interior and exterior solutions can be continuously connected for any k if the phase shifts are chosen according to equation (27). On the other hand, given an energy value E_{λ} , a bound state with this energy eigenvalue exists only when the potential $V_a(x)$ has a specific form such that the solutions in the two regions can be connected continuously. Here we are interested in the case of the critical energy $E = \mu \ (\lambda = 0)$. The exterior solutions are then given by (up to a normalization factor)

$$
u_{\mu}^{>}(x,\pm) = 1, \quad v_{\mu}^{>}(x,\pm) = 0.
$$
 (34)

They can be continuously connected to the interior ones only when $V_a(x)$ takes some specific form such that

$$
G_{\pm}(a,0) = 0.\t\t(35)
$$

On account of equation (33), we conclude that critical energy solutions with $E = \mu$ exist if and only if

$$
g_{\pm}(a,0) = 0.\t\t(36)
$$

From equation (34) we see that the critical energy solutions, if exist, are not bound states since they are not normalizable. Nevertheless, the argument below equation (18b) applies since $v_{\mu}^{\geq} = 0$. Thus the solution with $E = \mu$ is not degenerate. In other words, the two solutions with $E = \mu$ and different parities cannot appear simultaneously for a given potential. As in the free case these solutions are called half-bound states.

Now we easily realize that the first limit in equation (29) corresponds to the existence of critical energy states with $E = \mu$. It can then be verified that this corresponds to the first case in equations (30, 31). The other cases in these equations correspond to the case without the above critical energy states. Thus we conclude that $\eta_+(\mu)/\pi$ $|\eta_-(\mu)/\pi|$ takes integers (half integers) when these exists a critical state with $E = \mu$ and even(odd)parity, otherwise it takes half integers (integers). It is easy to check that $|T_0| = 1$ when there exists a half-bound state with $E = \mu$ (even or odd), otherwise $T_0 = 0$ (*cf.* Eq. (24)).

In the above we have analysed the threshold behaviour of $\eta_{\pm}(E_k)$ in detail. The threshold behaviour of $\eta_{\pm}(-E_k)$ can be discussed in a parallel way. The results are given by

$$
\tan \eta_{+}(-E_k) \to d_{+} \xi^{-(2q_{+}-1)} \quad \text{or} \quad \tilde{d}_{+} \xi^{2\tilde{q}_{+}-1}, \quad (\xi \to 0), \tag{37a}
$$

$$
\tan \eta_{-}(-E_k) \to d_{-} \xi^{2q_{-}-1}
$$
 or $\tilde{d}_{-} \xi^{-(2\tilde{q}_{-}-1)}$, $(\xi \to 0)$, (37b)

where q_{\pm} and \tilde{q}_{\pm} are natural numbers, d_{\pm} and \tilde{d}_{\pm} are nonzero constants. The first limit in equation (37a) or (37b) corresponds to the case when there exists a critical state with $E = -\mu$ and corresponding parity. Thus $\eta_{+}(-\mu)/\pi$ $[\eta_{-}(-\mu)/\pi]$ takes half integers (integers) when there exists a critical state with energy $E = -\mu$ and even(odd)-parity, otherwise it takes integers (half integers).

The critical energy states with $E = -\mu$, if exist, are given by (up to a normalization factor)

$$
u_{-\mu}^{>}(x,\pm) = 0, \quad v_{-\mu}^{>}(x,\pm) = 1 \tag{38}
$$

in the region $x>a$. These are also half-bound states. For a given potential, the two solutions with $E = -\mu$ cannot appears simultaneously.

Using the threshold behaviours obtained above it can be easily verified that when $k = 0$ equation (22) reduces to the above critical solutions or trivial solutions according as the corresponding critical solutions exist or not.

It may be beneficial to discuss some specific model and to verify the above threshold behaviour of the phase shifts. A typical and simple cutoff potential is the square well potential with depth V_0 and width $2a$. This model can be solved explicitly, though the results are far from simple. Here we are interested only in scattering states and critical states. The conditions for the existence of critical energy states can be worked out explicitly. For scattering states, one can find the closed form for $\tan \eta_{\pm}(E_k)$ and $\tan \eta_{+}(-E_k)$, and discuss the limit $k \to 0$. Since the calculations are somewhat tedious but straightforward, we will not give the details here. We just point out that the general results and conclusions obtained above are all confirmed by the simple model at hand.

4 The Levinson theorem

With the above preparations, we are now ready to establish the Levinson theorem. The theorem is developed on the basis of the completeness relations (21) and the threshold behavior of the phase shifts given in equations (31, 37). In addition, the fundamental equation (17) will be employed in the development of the theorem.

Equation (21) is a matrix equation of which the rhs is a diagonal matrix. When written in matrix elements, it gives four equations. For free particles, the last two terms on the lhs are absent. We write down the first diagonal equation. Replacing $x[′]$ by $-x[′]$ in this equation and using equation (18) we get another. Taking the sum

 $($

and the difference of these two equations we have

$$
\int_0^\infty dk \left[u_k^0(x, \pm) u_k^{0*}(x', \pm) + u_{-k}^0(x, \pm) u_{-k}^{0*}(x', \pm) \right]
$$

$$
= \frac{1}{2} [\delta(x - x') \pm \delta(x + x')], \quad (39)
$$

where the superscript 0 indicates free particles. Carrying out the same procedure to the second diagonal equation we have

$$
\int_0^\infty dk \left[v_k^0(x, \pm) v_k^{0*}(x', \pm) + v_{-k}^0(x, \pm) v_{-k}^{0*}(x', \pm) \right]
$$

$$
= \frac{1}{2} [\delta(x - x') \mp \delta(x + x')]. \quad (40)
$$

The sum of equations (39, 40) gives

$$
\int_0^\infty dk \sum_s [u_{ks}^0(x, \pm) u_{ks}^{0*}(x', \pm) + u_{-ks}^0(x, \pm) u_{-ks}^{0*}(x', \pm)] = \delta(x - x'), \quad (41)
$$

where the second notation in equation (16) has been used and the spinor index is denoted by s. A similar result for the case with an external symmetric potential reads

$$
\int_0^\infty dk \sum_s [u_{ks}(x, \pm) u_{ks}^*(x', \pm) + u_{-ks}(x, \pm) u_{-ks}^*(x', \pm)]
$$

$$
+ \sum_{\kappa_\pm} \sum_s u_{\kappa_\pm s}(x) u_{\kappa_\pm s}^*(x') = \delta(x - x'). \quad (42)
$$

Now we subtract equation (42) from equation (41) , then set $x' = x$, and integrate over x from $-\infty$ to $+\infty$, we arrive at

$$
n_{\pm} = \int_0^{\infty} dk \left[(\psi_k^0(\pm), \psi_k^0(\pm)) - (\psi_k(\pm), \psi_k(\pm)) \right] + \int_0^{\infty} dk \left[(\psi_{-k}^0(\pm), \psi_{-k}^0(\pm)) - (\psi_{-k}(\pm), \psi_{-k}(\pm)) \right],
$$
\n(43)

where we have used equation (20) to get

$$
\sum_{\kappa_\pm}(\psi_{\kappa_\pm},\psi_{\kappa_\pm})=\sum_{\kappa_\pm}1=n_\pm
$$

where n_{+} (n−) is the number of bound states with even(odd)-parity. The inner products are defined by integrals similar to that in equation (20). From orthonormal relations we have, say, $(\psi_k^0(\pm), \psi_k^0(\pm)) = \delta(0) \pm \delta(2k)$, and similarly for the other inner products in equation (43) . These are infinities and should be treated very carefully. In order to avoid the difficulty of infiniteness, we define

$$
(\psi_{k'}(\pm), \psi_k(\pm))_{r_0} \equiv \int_{-r_0}^{r_0} dx \, \psi_{k'}^{\dagger}(x, \pm) \psi_k(x, \pm), \quad (44)
$$

and obtain $(\psi_k(\pm), \psi_k(\pm))$ in the limit $k' \to k$ and $r_0 \rightarrow \infty$. The other inner products in equation (43) will be treated in the same way. Using equation (17) it can be shown that

$$
(\psi_{\pm k'}, \psi_{\pm k})_{r_0} = \pm \left. \frac{1}{E_{k'} - E_k} (v_{\pm k'}^* u_{\pm k} - u_{\pm k'}^* v_{\pm k}) \right|_{-r_0}^{r_0},
$$
\n(45)

which holds for both even-parity and odd-parity solutions. Using the asymptotic forms (22), and taking the limit $k' \rightarrow k$, we find

$$
(\psi_{\pm k}(+), \psi_{\pm k}(+))_{r_0} = \frac{r_0}{\pi} + \frac{1}{\pi} \frac{d\eta_{+}(\pm E_k)}{dk}
$$

$$
\pm \frac{\mu}{2\pi k E_k} \sin[2kr_0 + 2\eta_{+}(\pm E_k)],
$$
(46a)

$$
(\psi_{\pm k}(-), \psi_{\pm k}(-))_{r_0} = \frac{r_0}{\pi} + \frac{1}{\pi} \frac{d\eta_{-}(\pm E_k)}{dk}
$$

$$
= \frac{\mu}{2\pi k E_k} \sin[2kr_0 + 2\eta_{-}(\pm E_k)].
$$
(46b)

For free particles, the corresponding results are obtained by setting the phase shifts to zero in the above equations. Obviously, the infiniteness in these results lies in the first term r_0/π in each equation when the limit $r_0 \to \infty$ is taken. This disappears when we take the difference of equation (46) and the corresponding results for free particles. Using the well-known formulas

$$
\lim_{r_0 \to \infty} \frac{\sin 2kr_0}{\pi k} = \delta(k),
$$

and $g(k)\delta(k) = g(0)\delta(k)$ for any continuous function $g(k)$, we obtain

$$
(\psi_{\pm k}^{0}(+), \psi_{\pm k}^{0}(+))_{r_{0}} - (\psi_{\pm k}(+), \psi_{\pm k}(+))_{r_{0}} = -\frac{1}{\pi} \frac{d\eta_{+}(\pm E_{k})}{dk}
$$

$$
\pm \sin^{2} \eta_{+}(\pm \mu) \delta(k) \mp \frac{\mu}{2\pi k E_{k}} \cos 2kr_{0} \sin 2\eta_{+}(\pm E_{k}),
$$
(47a)

$$
\psi_{\pm k}^{0}(-), \psi_{\pm k}^{0}(-)\rangle_{r_{0}} - (\psi_{\pm k}(-), \psi_{\pm k}(-)\rangle_{r_{0}} = -\frac{1}{\pi} \frac{d\eta_{-}(\pm E_{k})}{dk}
$$

$$
\mp \sin^{2} \eta_{-}(\pm \mu) \delta(k) \pm \frac{\mu}{2\pi k E_{k}} \cos 2kr_{0} \sin 2\eta_{-}(\pm E_{k}).
$$
\n(47b)

We regroup the four equations contained in equation (47) according to the sign of the energy instead of the parity, then integrate each group over k from 0 to $+\infty$, and take the limit $r_0 \rightarrow \infty$, we have

$$
\int_0^\infty dk \left[(\psi_k^0(\pm), \psi_k^0(\pm)) - (\psi_k(\pm), \psi_k(\pm)) \right]
$$

= $\frac{1}{\pi} [\eta_{\pm}(\mu) - \eta_{\pm}(+\infty)] \pm \frac{1}{2} \sin^2 \eta_{\pm}(\mu)$
 $\mp \frac{\mu}{2\pi} \lim_{r_0 \to \infty} \int_0^\infty dk \frac{\sin 2\eta_{\pm}(E_k)}{kE_k} \cos 2kr_0$, (48a)

$$
\int_0^\infty dk \left[(\psi_{-k}^0(\pm), \psi_{-k}^0(\pm)) - (\psi_{-k}(\pm), \psi_{-k}(\pm)) \right]
$$

= $\frac{1}{\pi} [\eta_{\pm}(-\mu) - \eta_{\pm}(-\infty)] \mp \frac{1}{2} \sin^2 \eta_{\pm}(-\mu)$
 $\pm \frac{\mu}{2\pi} \lim_{r_0 \to \infty} \int_0^\infty dk \frac{\sin 2\eta_{\pm}(-E_k)}{kE_k} \cos 2kr_0,$ (48b)

where we have used the integral $\int_0^\infty dk \, \delta(k) =$ $(1/2)$ $\int_{-\infty}^{+\infty} dk \, \delta(k) = 1/2$ since the Dirac δ function is an even function. So far in this section we have not cut off the potential. In the following we set $V(x) = V_a(x)$. Then we have the threshold behavior (Eqs. (31, 37)). The last term in equation (48a) can be decomposed into two integrals, the first from 0 to $\varepsilon = 0^+$ while the second from ε to $+\infty$. The second integral vanishes in the limit $r_0 \rightarrow \infty$ since the factor $\cos 2kr_0$ oscillates very rapidly and the other factors in the integrand are finite. For the first integral, we have

$$
\int_0^{\varepsilon} dk \frac{\sin 2\eta_{\pm}(E_k)}{kE_k} \cos 2kr_0 = \frac{1}{\mu} \int_0^{\varepsilon} dk \frac{\sin 2\eta_{\pm}(E_k)}{k}
$$

$$
= \frac{1}{\mu} \int_0^{\varepsilon a} d\xi \frac{\sin 2\eta_{\pm}(E_k)}{\xi}, \tag{49}
$$

since $k \leq \varepsilon$ is very small. On account of equation (31), we have $\sin 2\eta_{\pm}(E_k) \rightarrow c_{\pm} \xi^{2r_{\pm}-1}$ ($\xi \rightarrow 0$), where r_{\pm} are natural numbers and c_{\pm} are nonzero constants. Substituting into the above equation we have

$$
\int_0^{\varepsilon a} d\xi \, \frac{\sin 2\eta_{\pm}(E_k)}{\xi} = \frac{c_{\pm}(\varepsilon a)^{2r_{\pm}-1}}{2r_{\pm}-1} \to 0,
$$
\n
$$
(\varepsilon \to 0^+).
$$
\n(50)

Thus the first integral vanishes as well. Then the last term in equation (48a) vanishes. By using equation (37) it can be shown in a similar way that the last term in equation (48b) also vanish. Substituting the results (Eqs. (48a, 48b)), each without the last term, into equation (43) we arrive at

$$
[\eta_{\pm}(\mu) - \eta_{\pm}(+\infty)] + [\eta_{\pm}(-\mu) - \eta_{\pm}(-\infty)]
$$

$$
\pm \frac{\pi}{2} [\sin^2 \eta_{\pm}(\mu) - \sin^2 \eta_{\pm}(-\mu)] = n_{\pm} \pi. \quad (51)
$$

This is the Levinson theorem for Dirac particles in an external symmetric potential $V_a(x)$ in one dimension. It relates the phase shifts to the total number of bound states for each parity. In the next section we discuss some relevant problems and study two examples. Finally we summarize the results briefly.

5 Discussions

5.1 Phase shifts at infinite momentum

It should be pointed out that there is no modulo- π ambiguity in equation (51), because only difference of phase shifts at different momentums and trigonometric functions of the phase shifts are involved. However, it may be beneficial to appropriately define the phase shifts such that they can be determined uniquely. We consider the potential $\theta V(x)$ where $0 \le \theta \le 1$ is a parameter independent of x, and $V(x)$ is symmetric but not necessarily be $V_a(x)$. Positive-energy scattering solutions in this potential will be denoted by $\psi_k(x, \pm, \theta)$ and negative-energy ones by $\psi_{-k}(x, \pm, \theta)$. The corresponding phase shifts will be denoted by $\eta_{\pm}(E_k, \theta)$ and $\eta_{\pm}(-E_k, \theta)$ respectively. It is natural to define

$$
\eta_{\pm}(E_k, 0) = 0, \quad \eta_{\pm}(-E_k, 0) = 0, \tag{52}
$$

since there is no potential in this case. We also require that $\eta_{\pm}(E_k, \theta)$ and $\eta_{\pm}(-E_k, \theta)$ be continuous functions of θ for any finite k. Then the phase shifts $\eta_{\pm}(E_k) = \eta_{\pm}(E_k, 1)$, $\eta_{\pm}(-E_k) = \eta_{\pm}(-E_k, 1)$ in the potential $V(x)$ are definitely defined. It should be remarked that the phase shifts at threshold are not continuous in θ , however (this was discussed in some detail in Ref. [18]).

To determine the phase shifts at infinite momentum, we use equations (17, 22) for two potentials $\theta V(x)$ and $\hat{\theta}V(x)$. It can be shown that

$$
\sin[\Delta \eta_+(\pm E_k, \theta)] = \mp \Delta \theta \frac{\pi E_k}{k} (\psi_{\pm k}(+,\tilde{\theta}), V\psi_{\pm k}(+,\theta)),
$$
\n(53)

where $\Delta\theta = \tilde{\theta} - \theta$ and

$$
\Delta \eta_+(\pm E_k, \theta) = \eta_+(\pm E_k, \tilde{\theta}) - \eta_+(\pm E_k, \theta).
$$

In the limit $\tilde{\theta} \to \theta$ or $\Delta \theta \to 0$, the $\psi_{+k}(x, +, \tilde{\theta})$ in equation (53) can be replaced by $\psi_{\pm k}(x, +, \theta)$, and the sine on the lhs can be replaced by its argument since the phase shifts are continuous in θ as required above, so we have

$$
\frac{\mathrm{d}\eta_+(\pm E_k,\theta)}{\mathrm{d}\theta} = \mp \frac{\pi E_k}{k} (\psi_{\pm k}(+,\theta), V\psi_{\pm k}(+,\theta)). \tag{54}
$$

Integrating over θ from 0 to 1 and using equation (52) we have

$$
\eta_{+}(\pm E_k) = \mp \frac{\pi E_k}{k} \int_0^1 d\theta \, (\psi_{\pm k}(+,\theta), V\psi_{\pm k}(+,\theta)). \tag{55}
$$

For $\eta_-(\pm E_k)$ we have a similar result. These results are practically not useful since the solutions on the rhs are not explicitly available in general. However, when $k \to \infty$, we can ignore the potential $\theta V(x)$ in equation (17) since $V(x)$ is not very singular at $x = 0$ (less singular than x^{-1}) as assumed) and is regular elsewhere. Then the solutions $\psi_{\pm k}(x, +, \theta)$ in equation (55) can be replaced by the free ones $\psi_{\pm k}^0(x,+)$ and we have

$$
\eta_+(\pm \infty) = \mp \int_0^\infty dx \, V(x). \tag{56}
$$

The result for $\eta_-(\pm\infty)$ is the same. The integral in the above equation converges because $V(x)$ decreases more rapidly than x^{-2} when $x \to \infty$ and is less singular than x^{-1} when $x \to 0$, as assumed in Section 2. These results are similar to those obtained in three dimensions [9,16,17] and two dimensions [18]. Of course, they hold in the special case $V(x) = V_a(x)$. As a consequence, we have

$$
\eta_{\pm}(+\infty)+\eta_{\pm}(-\infty)=0, \qquad (57)
$$

and the Levinson theorem (51) reduces to the form of equation (7).

We have calculated $\eta_+(\pm\infty)$ and $\eta_-(\pm\infty)$ exactly for the square well potential, and the results (56) are confirmed. In the following we will see that equation (56) does not hold for the δ potential well, however. This is because the δ potential well is more singular than the type we have assumed. Indeed, if $V(x)$ is less singular than x^{-1} at $x = 0$, or behaves like $A_0 x^{-1+\rho/2}$, we have $\int_0^{\varepsilon} dx V(x) =$ $2A_0\varepsilon^{\rho/2}/\rho \to 0 \ (\varepsilon \to 0^+).$ If $V(x) = -U_0\delta(x)$, however, we have $\int_0^{\varepsilon} dx V(x) = -U_0/2$. Thus the δ potential well is more singular. It is expected that equation (56) remains correct as long as $V(x)$ belongs to the type we assumed, *i.e.*, less singular than x^{-1} at $x = 0$.

5.2 Verification of the theorem

To examine the Levinson theorem one should choose a simple potential such that n_{\pm} and all phase shifts at zero momentum can be worked out explicitly. This is not available even for the square well potential. Although closed forms for $\tan \eta_{\pm}(E_k)$ and $\tan \eta_{\pm}(-E_k)$ can be obtained, they depend on k in a very complicated way. To determine $\eta_+(\mu)$ and $\eta_+(-\mu)$, numerical calculations are necessary. The transcendental equations for the energy levels of bound states are also complicated. Perhaps the simplest potential is the δ potential well $V(x) = -U_0 \delta(x)$, where $U_0 > 0$ is a dimensionless parameter. We have pointed out above that this potential does not belong to the type we have assumed in developing the Levinson theorem. However, threshold behavior of the phase shifts similar to equations (31, 37) can be explicitly shown. So the Levinson theorem should remain correct in this case. As the calculations are simple, we only give the results. We have $n_{+} = 1, n_{-} = 0$. The phase shifts at infinite momentum are different from those given by equation (56):

$$
\eta_+(\pm \infty) = \pm \arctan \frac{U_0}{2},\tag{58}
$$

where $\arctan(U_0/2) \in (0, \pi/2)$ is the principal value. A similar result holds for $\eta_-(\pm\infty)$. Thus equation (57) remains valid in this case though equation (56) does not, and the reduced form of the Levinson theorem (7) is expected to hold. In fact, the phase shifts at threshold can be found to be

$$
\eta_{+}(\mu) = \frac{\pi}{2}, \quad \eta_{+}(-\mu) = 0,
$$

$$
\eta_{-}(\mu) = 0, \quad \eta_{-}(-\mu) = -\frac{\pi}{2},
$$
 (59)

and it is easy to verify that equation (7) is satisfied.

For a δ potential barrier $V(x) = U_0 \delta(x)$ where $U_0 > 0$, it can be shown that $n_{+} = 0$, $n_{-} = 1$, $\eta_{+}(\mu) = -\pi/2$, $\eta_{+}(-\mu) = 0$, $\eta_{-}(\mu) = 0$, $\eta_{-}(-\mu) = \pi/2$, and equation (7) is satisfied as well.

A less simple example is the double δ potential wells $V(x) = -U_0[\delta(x-a) + \delta(x+a)]$, where $U_0 > 0$ is dimensionless. This example is somewhat more substantial since half-bound states may be involved, and both n_{\pm} and the phase shifts depend on the value of U_0 . Since the calculations are not difficult but lengthy, we will not give the details. We only point out that in this case the Levinson theorem is confirmed once again.

5.3 Summary

In this paper we study Dirac particles in one-dimensional symmetric potentials that decrease more rapidly than x^{-2} when $x \to \infty$ and are less singular than x^{-1} when $x \to 0$. The properties of bound-state and scattering solutions are discussed. For cutoff potentials the threshold behaviours of the phase shifts are studied in detail, and the Levinson theorem (7) is established, which connects the phase shifts with the total number of bound states for each parity. Two simple examples are discussed and the Levinson theorem is verified explicitly. A mathematically rigorous extension of the Levinson theorem to non-cutoff potentials is still not available. However, one may expect that the theorem remains correct for short-range non-cutoff potentials where the asymptotic forms (24) for scattering solutions hold and the total number of bound states is finite.

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